

# **An Overview of Fluid Animation**

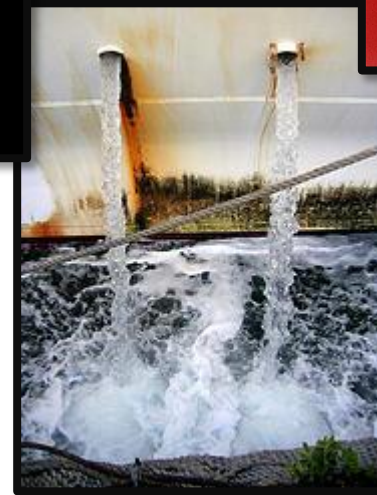
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Jan 13, 2016

# Reminders

- 1<sup>st</sup> round presentations start Monday.  
Graded on...
  1. Knowledge/coverage of technical concepts
  2. Organization
  3. Slide quality
  4. Speaking/presentation skills
- 1<sup>st</sup> round of paper reviews due Sunday, 5pm.
- Start thinking about project topics.
- Piazza (or email) with any questions.

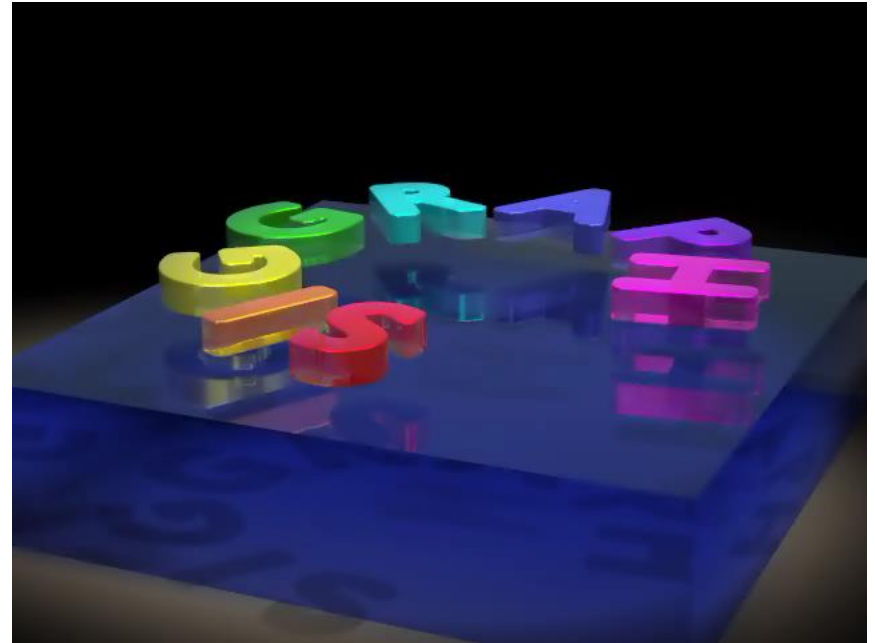
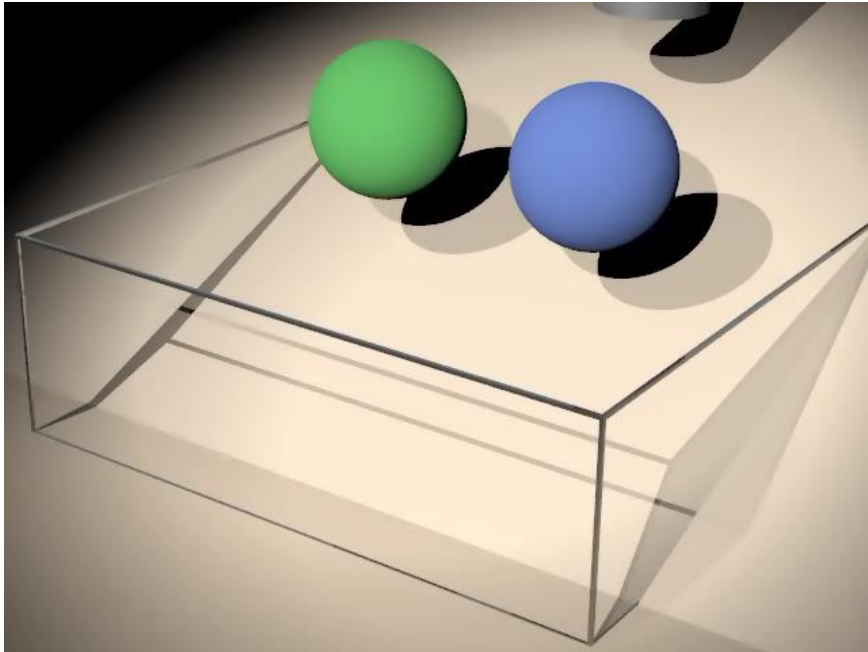
# What distinguishes fluids?



# What distinguishes fluids?

- No “preferred” shape.
- Always flows when force is applied.
- Deforms to fit its container.
- Internal forces depend on *velocities*, not displacements/deformation (compare w/ elastic objects)

# Examples



For further detail on today's material, see Robert Bridson's online fluid notes.

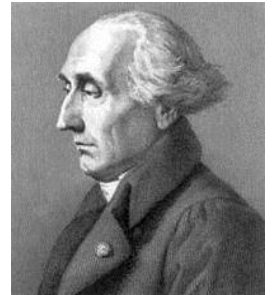
<http://www.cs.ubc.ca/~rbridson/fluidsimulation/>

(There's also a book, which is available in the library.)

# Basic Theory



# Eulerian vs. Lagrangian



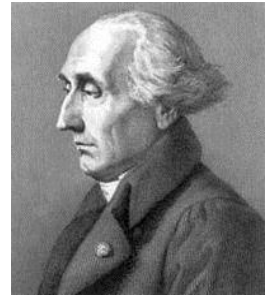
Lagrangian: Point of reference moves *with* the material.

Eulerian: Point of reference is *stationary*.

e.g. Weather balloon (Lagrangian) vs. weather station on the ground (Eulerian)

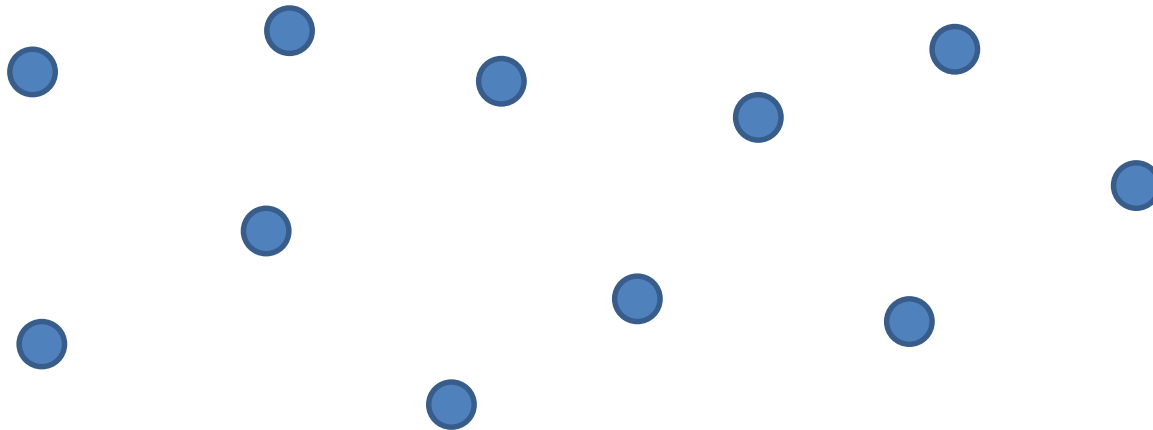


# Eulerian vs. Lagrangian



Consider an evolving scalar field (e.g., temperature).

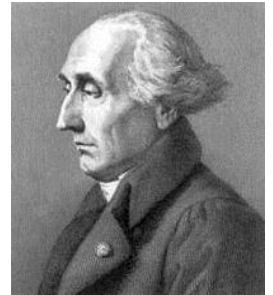
Lagrangian view: Set of *moving particles*, each with a temperature value.





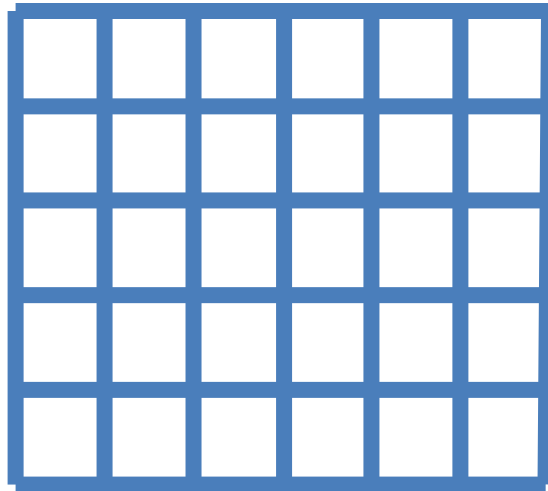


# Eulerian vs. Lagrangian



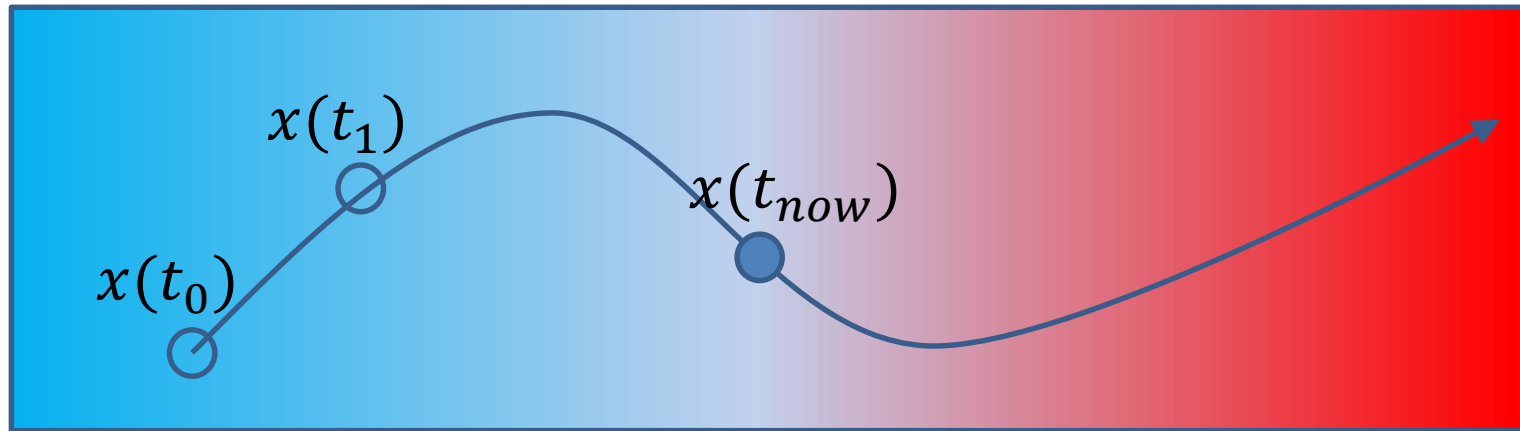
Consider an evolving scalar field (e.g., temperature).

Eulerian view: A fixed grid of temperature values, that temperature *flows through*.



# Relating Eulerian and Lagrangian

Consider the temperature  $T(x, t)$  at a point following a given path,  $x(t)$ .



How can the temperature measured at  $x(t)$  change?

1. There is a hot/cold “source” at the current point.
2. Following the path, the point moves to a cooler/warmer location.

# Time derivatives

Mathematically:

$$\frac{D}{Dt} T(x(t), t) = \frac{\partial T}{\partial t} + \frac{\partial T}{\partial x} \frac{\partial x}{\partial t}$$

Chain rule!

$$= \frac{\partial T}{\partial t} + \nabla T \cdot \frac{\partial x}{\partial t}$$

Definition  
of  $\nabla$

$$= \frac{\partial T}{\partial t} + \mathbf{u} \cdot \nabla T$$

Choose  
 $\frac{\partial x}{\partial t} = \mathbf{u}$

# Material Derivative

This is called the *material derivative*, and denoted  $\frac{D}{Dt}$ .  
(AKA total derivative.)

Change at a point moving  
along the given path,  $x(t)$ .

Change due  
to movement  
of the point.

$$\frac{DT}{Dt} = \frac{\partial T}{\partial t} + \mathbf{u} \cdot \nabla T$$

Change at the current  
(fixed) point.

# Advection

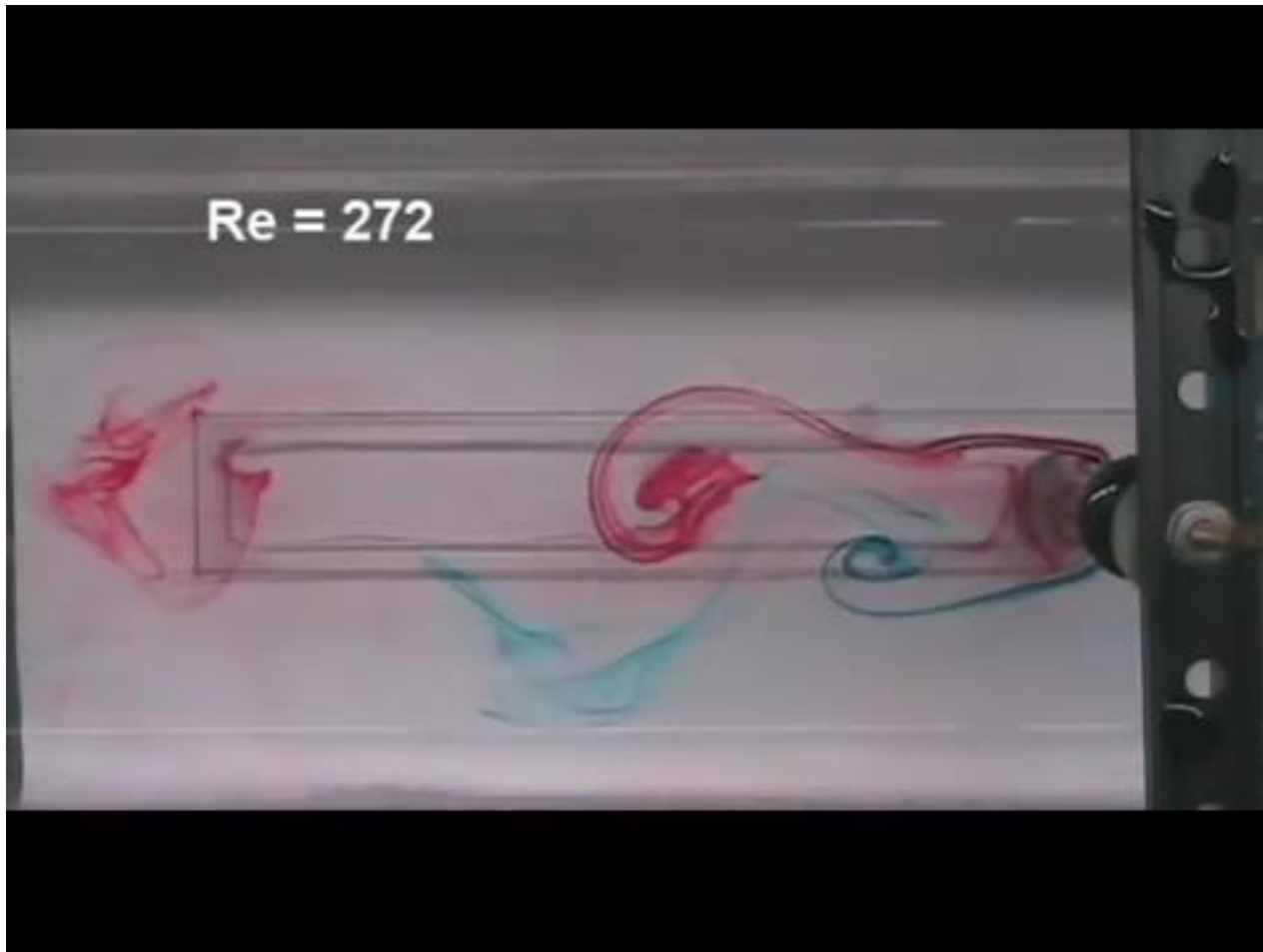
To track a quantity  $T$  moving (passively) through a velocity field:

$$\frac{DT}{Dt} = 0 \quad \text{or equivalently} \quad \frac{\partial T}{\partial t} + \mathbf{u} \cdot \nabla T = 0$$

This is the *advection equation*.

Think of colored dye or massless particles drifting around in fluid.

# Advection



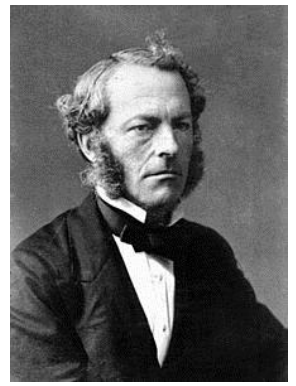
# Equations of Motion

For general continuum materials, we essentially had Newton's second law:  $F = ma$ .

The *Navier-Stokes equations* are the same equations again, specialized to fluids.



# Navier-Stokes



Density  $\times$  Acceleration = Sum of Forces

$$\rho \frac{D\mathbf{u}}{Dt} = \sum_i \mathbf{F}_i$$

Expanding the material derivative...

$$\rho \frac{\partial \mathbf{u}}{\partial t} = -\rho(\mathbf{u} \cdot \nabla \mathbf{u}) + \sum_i \mathbf{F}_i$$



# What are the forces on a fluid?

Primarily:

- Pressure
- Viscosity
- Simple “external” forces
  - (e.g. gravity, buoyancy, user/artistic forces)

Also:

- Surface tension
- Coriolis
- Possibilities for more exotic fluid types:
  - Elasticity (e.g. silly putty)
  - Shear thickening / thinning (e.g. “oobleck”, ketchup, paints)
  - Electromagnetic forces: magnetohydrodynamics, ferrofluids, etc.
- Various others...

# Exotic Fluids - Oobleck



# Oobleck Simulation

## Oobleck: Viscoplastic v.s. Shear-Thickening

Simulation parameters (viscoplastic):

$\rho$ : 1000.0 kg/m<sup>3</sup>  $\kappa$ : 109.0 kPa  $\mu$ : 11.2 kPa  $\sigma_Y$ : 0.1 Pa  $\eta$ : 10.0  $m$ : 1.0  $\sigma_T$ : 1.0  $\eta_p$ : 0.3  
#points: 519171-534871 grid res.: 157×157×157 dt: 0.5×10<sup>-5</sup> s subgrid geom. rem.: no

Simulation parameters (shear-thickening):

$\rho$ : 1000.0 kg/m<sup>3</sup>  $\kappa$ : 109.0 kPa  $\mu$ : 11.2 kPa  $\sigma_Y$ : 0.1 Pa  $\eta$ : 10.0  $m$ : 2.8  $\sigma_T$ : 1.0  $\eta_p$ : 0.3  
#points: 519171-529365 grid res.: 157×157×157 dt: 0.5×10<sup>-5</sup> s subgrid geom. rem.: no

[Yue et al. 2015]

# Exotic Fluids - Ferrofluid



# Fluid equations of motion...

$$\rho \frac{\partial \mathbf{u}}{\partial t} = -\rho(\mathbf{u} \cdot \nabla \mathbf{u}) + \sum_i \mathbf{F}_i$$

Change in  
velocity at a  
*fixed point*

Advection (of  
*velocity*)

Forces  
(pressure,  
viscosity  
gravity,...)

# Operator splitting

Break the full, nonlinear equation into sub-steps:

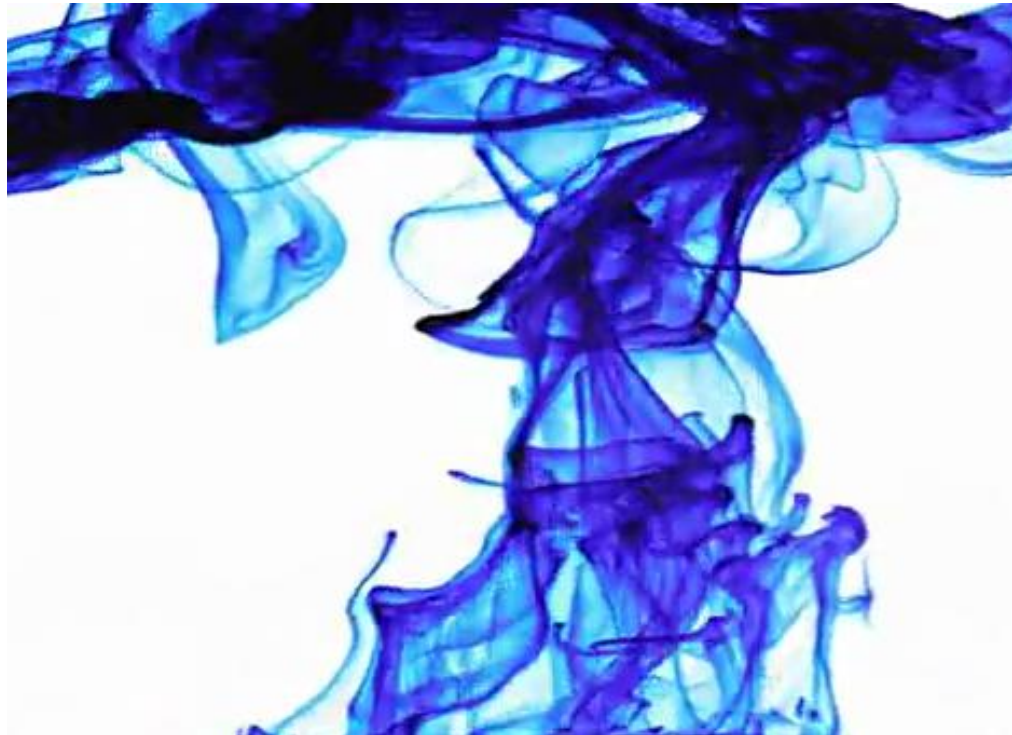
1. Advection:  $\rho \frac{\partial \mathbf{u}}{\partial t} = -\rho(\mathbf{u} \cdot \nabla \mathbf{u})$

2. Pressure:  $\rho \frac{\partial \mathbf{u}}{\partial t} = \mathbf{F}_{pressure}$

3. Viscosity:  $\rho \frac{\partial \mathbf{u}}{\partial t} = \mathbf{F}_{viscosity}$

4. External:  $\rho \frac{\partial \mathbf{u}}{\partial t} = \mathbf{F}_{other}$

# 1. Advection



# Advection

We already considered advection of a passive scalar quantity,  $T$ , under velocity  $\mathbf{u}$ .

$$\frac{\partial T}{\partial t} = -\mathbf{u} \cdot \nabla T$$

In Navier-Stokes advection term, we have:

$$\frac{\partial \mathbf{u}}{\partial t} = -\mathbf{u} \cdot \nabla \mathbf{u}$$

Velocity  $\mathbf{u}$  is advected (carried along) *by itself*, too!



# Advection

That is,  $(u, v, w)$  components of velocity  $\mathbf{u}$  are advected as separate scalars.

Can often reuse the same numerical method.

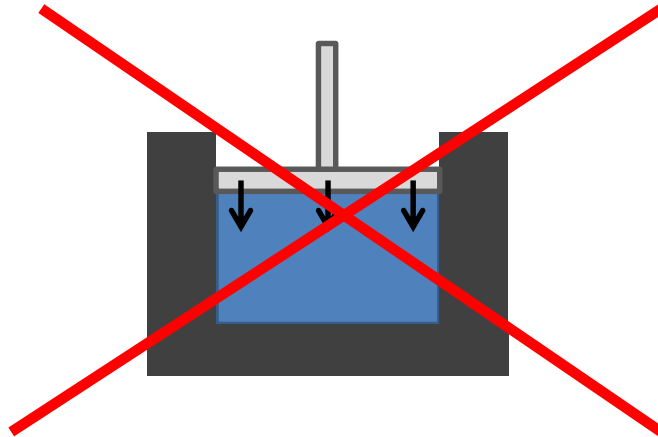
# 2. Pressure



# Pressure

What does pressure do?

- Enforces *incompressibility* (fights compression).



Typical fluids (mostly) do not *visibly* compress.

- Exceptions: high velocity, high pressure, ...

# Incompressibility

Compressible  
velocity field



Incompressible  
velocity field



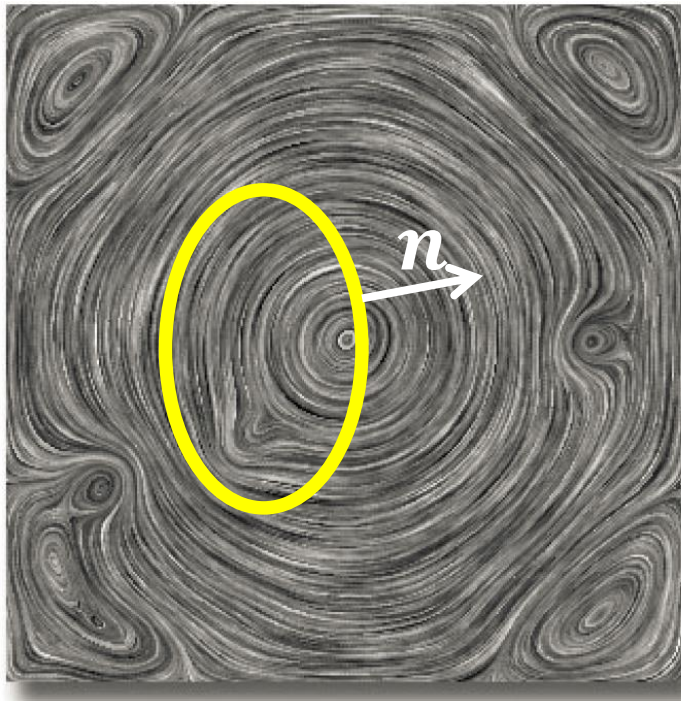


# Incompressibility

Intuitively, net flow into/out of a given region should be zero (no sinks/sources).

Integrate the net flow across the boundary of a closed region (yellow):

$$\int_{\partial\Omega} \mathbf{u} \cdot \mathbf{n} = 0$$



# Incompressibility

$$\int_{\partial\Omega} \mathbf{u} \cdot \mathbf{n} = 0$$

By divergence theorem:

$$\iint_{\Omega} \nabla \cdot \mathbf{u} = 0$$

But this is true for *any* region, so  $\nabla \cdot \mathbf{u} = \mathbf{0}$  everywhere.

Incompressibility implies  $\mathbf{u}$  is *divergence-free*.

# Pressure

Where does pressure come in?

- Pressure is the force needed to ensure the incompressibility constraint,  $\nabla \cdot \mathbf{u} = 0$ .
- Pressure force has the following form:

$$\mathbf{F}_p = -\nabla p$$

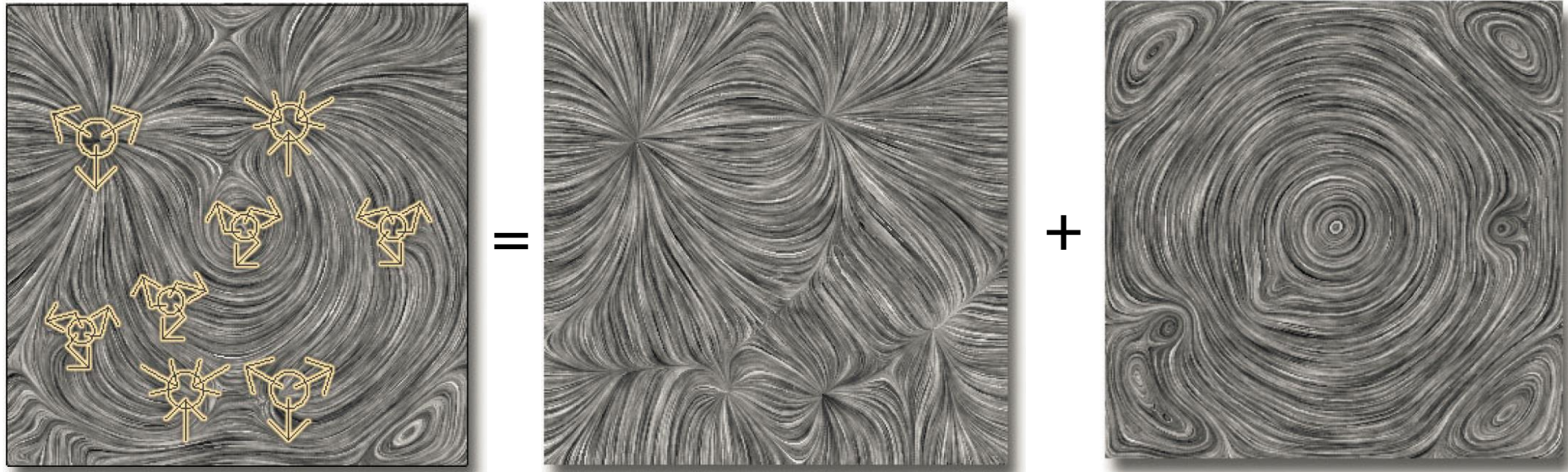
Let's see why...

# Helmholtz Decomposition

Input (Arbitrary)  
Velocity Field

Curl-Free  
(Irrotational)

Divergence-Free  
(Incompressible)



$$u = \nabla p + \nabla \times \varphi$$

$$u_{old} = F_{pressure} + u_{new}$$



# Aside: Pressure as Lagrange Multiplier

Interpret as an optimization:

Find the closest  $\mathbf{u}_{new}$  to  $\mathbf{u}_{old}$  where  $\nabla \cdot \mathbf{u}_{new} = 0$

$$\begin{aligned} & \underset{\mathbf{u}_{new}}{\operatorname{argmin}} \frac{\rho}{2} \|\mathbf{u}_{new} - \mathbf{u}_{old}\|^2 \\ & \text{subject to } \nabla \cdot \mathbf{u}_{new} = 0 \end{aligned}$$

The *Lagrange multiplier* for the incompressibility constraint is the pressure.

# 3. Viscosity



# High Speed Honey



# Viscosity

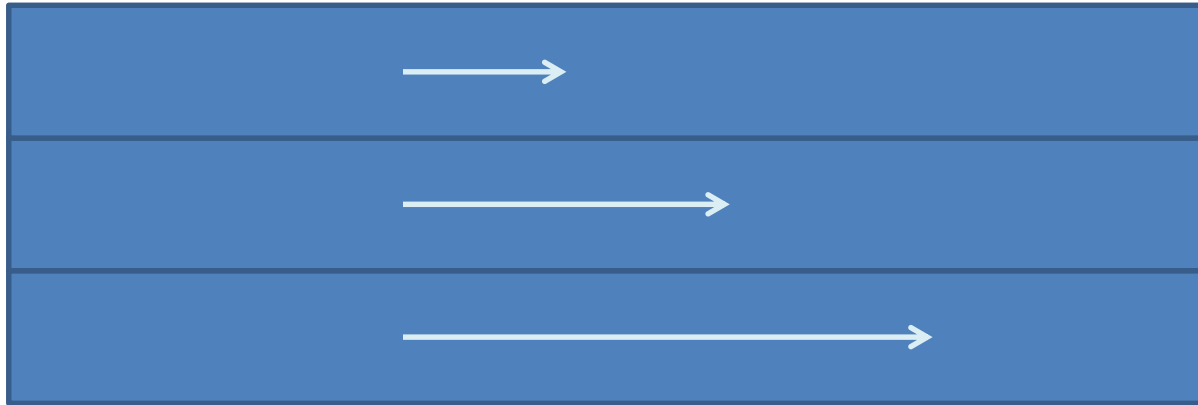


What characterizes a viscous liquid?

- “Thick”, highly *damped* behaviour.
- Strong resistance to flow.

# Viscosity

Loss of energy due to internal friction between molecules moving at different velocities.

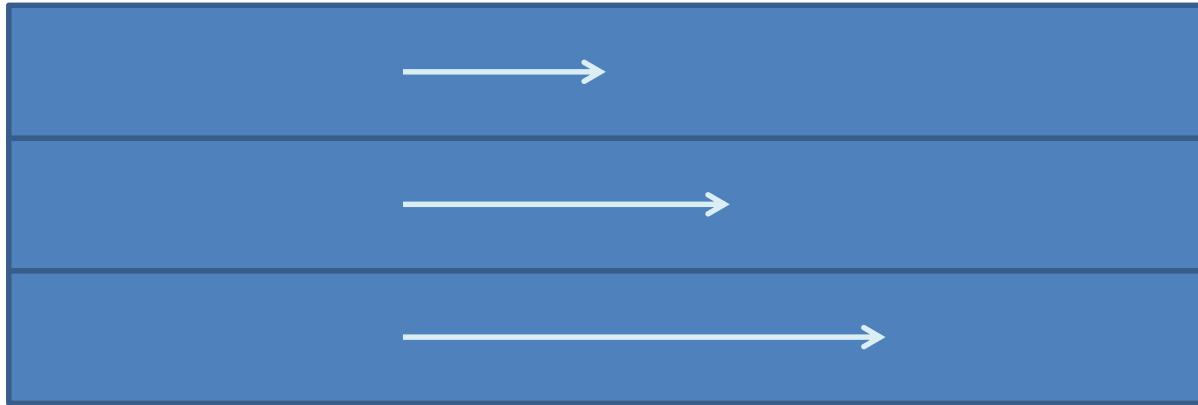


Interactions between molecules causes *shear stress* that...

- opposes *relative* motion.
- causes an exchange of momentum.

# Viscosity

Loss of energy due to internal friction between molecules moving at different velocities.

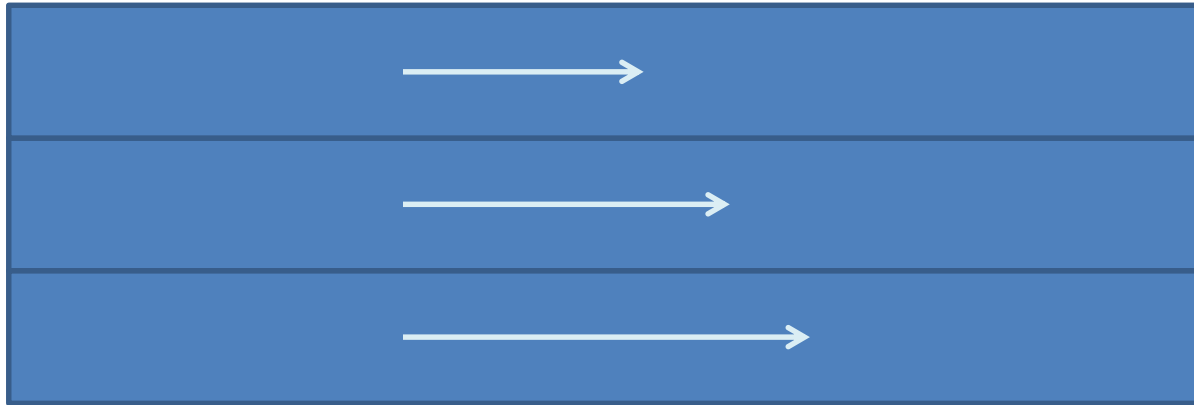


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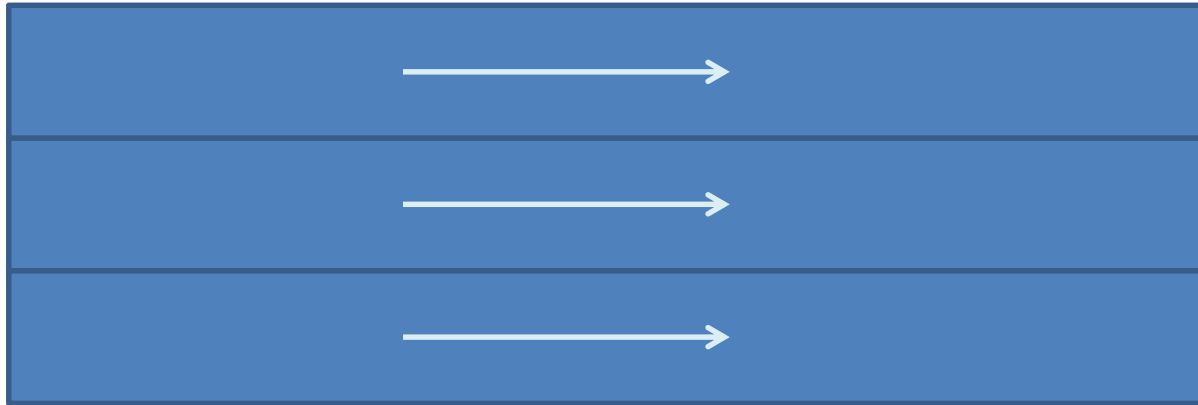


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# Viscosity

Loss of energy due to internal friction between molecules moving at different velocities.



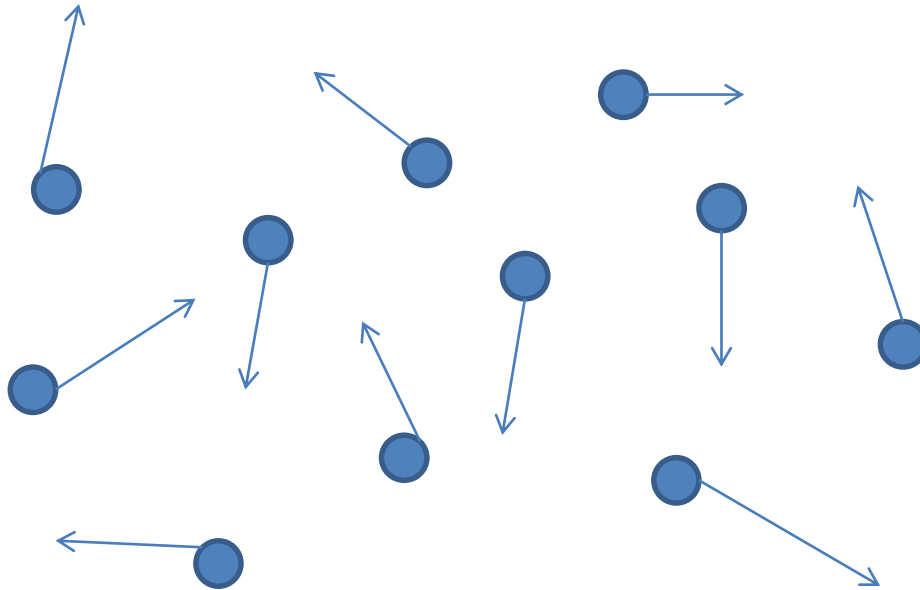
Interactions between molecules causes *shear stress* that...

- opposes *relative* motion.
- causes an exchange of momentum.



# Viscosity

Imagine fluid particles with general velocities.



Each particle interacts with nearby neighbours, exchanging momentum. Velocities gradually tend towards uniformity.

# Diffusion

The momentum exchange is related to:

- Velocity gradient,  $\nabla \mathbf{u}$ , in a region.
- Viscosity coefficient,  $\mu$ .

Net effect is a smoothing or *diffusion* of the velocity over time.

# Viscosity

Diffusion is typically modeled using the heat equation:

$$\frac{\partial T}{\partial t} = \alpha \nabla \cdot \nabla T$$

(e.g. modeling dye or heat spreading through a region.)



**Diffusion**

# Viscosity

Diffusion applied to velocity gives our viscous force:

$$\mathbf{F}_{viscosity} = \rho \frac{\partial \mathbf{u}}{\partial t} = \mu \nabla \cdot \nabla \mathbf{u}$$

Usually, we can diffuse each scalar component of the vector field  $\mathbf{u} = (u, v, w)$  separately.

# 4. External Forces



## Gravity.

It's not just a good idea.  
It's the Law.



# External Forces

Any other forces you may want.

- Simplest is gravity:
  - $F_g = \rho \mathbf{g}$  for  $\mathbf{g} = (0, -9.81, 0)$
- Buoyancy models are similar,
  - e.g.,  $F_b = \beta(T_{current} - T_{ref})\mathbf{g}$
- Artistic/user controls

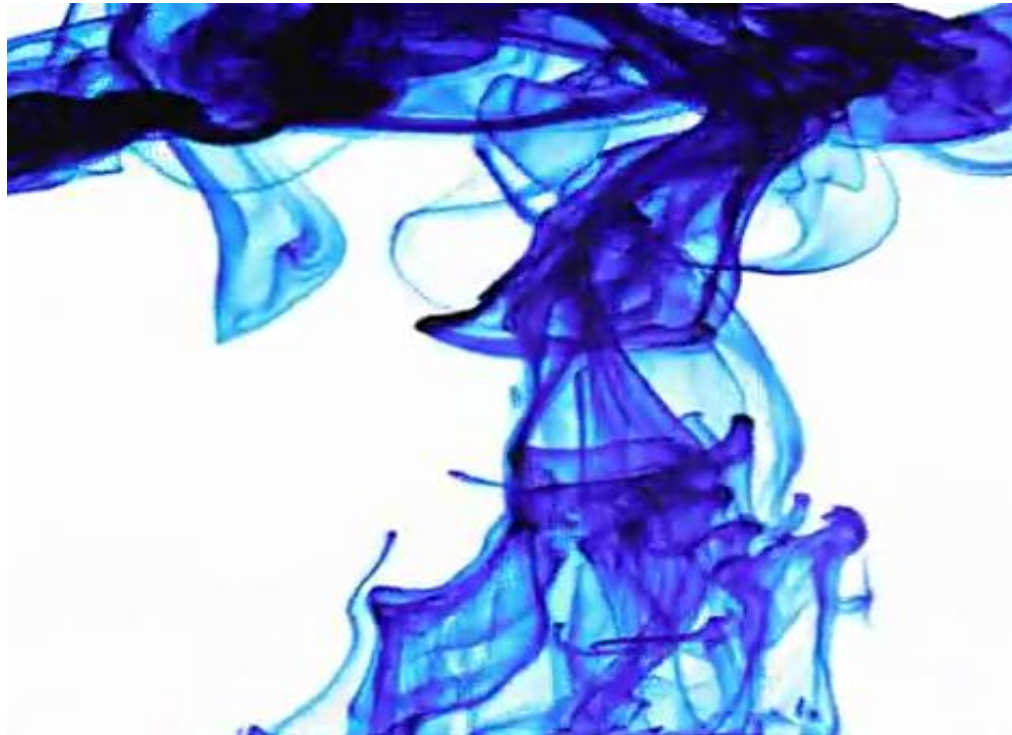
**e.g. Artistic control – Liquid Monster**



# **Discretizing the Equations for Fluid Animation**



# 1. Advection



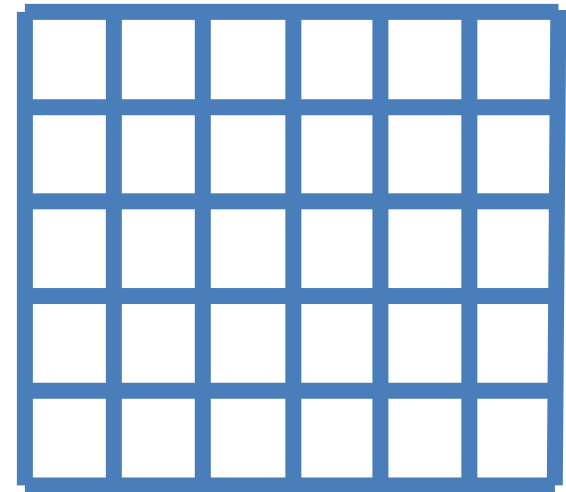
# Advection of a Scalar

Consider advecting a quantity,  $\varphi$

– temperature, color, smoke density, ...

according to a (fixed) velocity field  $\mathbf{u}$ .

Allocate a grid (2D array) that stores scalar  $\varphi$  and velocity  $\mathbf{u}$ .



# Eulerian

Approximate derivatives with *finite differences*.

$$\frac{\partial \varphi}{\partial t} + \mathbf{u} \cdot \nabla \varphi = 0$$

FTCS = Forward Time, Centered Space:

$$\frac{\varphi_i^{n+1} - \varphi_i^n}{\Delta t} + u \frac{\varphi_{i+1}^n - \varphi_{i-1}^n}{2\Delta x} = 0$$

Unconditionally  
Unstable!

Lax:

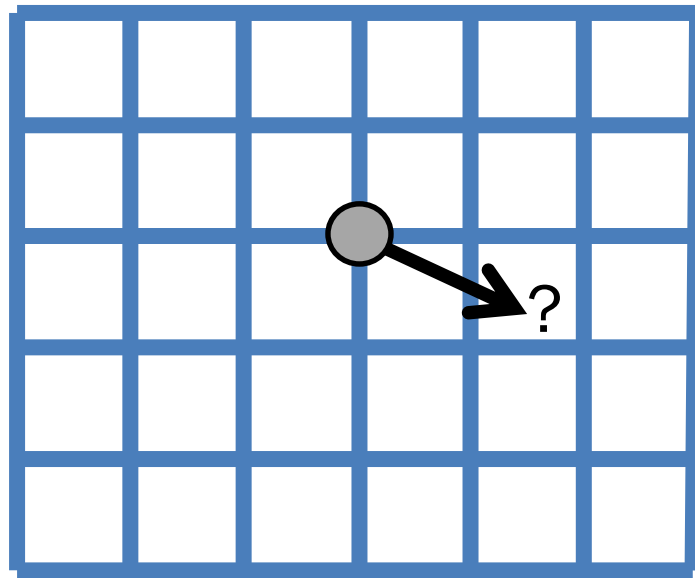
$$\frac{\varphi_i^{n+1} - (\varphi_{i+1}^n + \varphi_{i-1}^n)/2}{\Delta t} + u \frac{\varphi_{i+1}^n - \varphi_{i-1}^n}{2\Delta x} = 0$$

Conditionally  
Stable!

Many possible methods, stability can be a challenge.

# Lagrangian

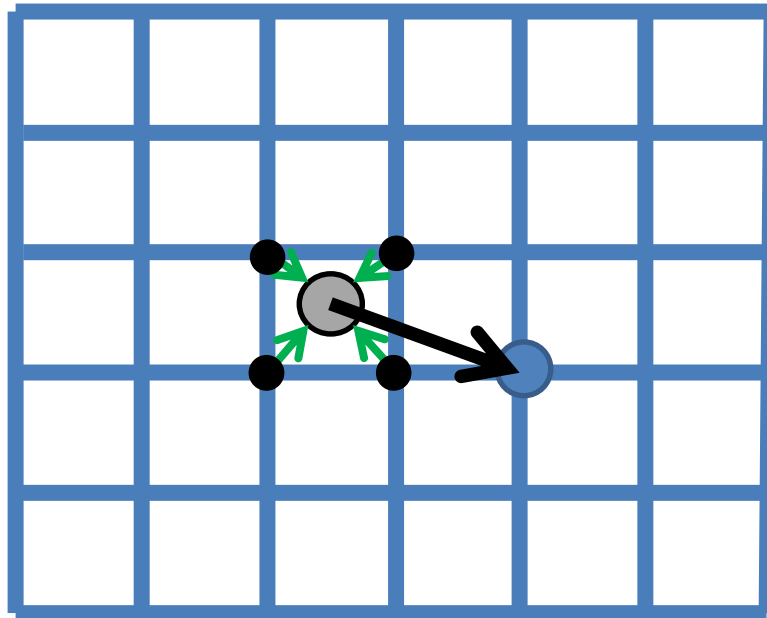
Advect data “forward” from grid points by integrating position according to grid velocity (e.g. forward Euler).



Problem: New data position doesn't necessarily land back on a grid point.

# Semi-Lagrangian

- Look *backwards* in time from a grid point (blue), to see where its new data is coming *from* (black).
- Interpolate data at the previous time position.



# Semi-Lagrangian - Details

1. Look up velocity  $\mathbf{u}_{i,j}$  at grid point.
2. Integrate position for a timestep of  $-\Delta t$  (FE).
  - e.g.  $x_{back} = x_{i,j} - \Delta t \mathbf{u}_{i,j}$
3. (Bilinearly) Interpolate  $\varphi$  at  $x_{back}$ , call it  $\varphi_{back}$ .
4. Assign  $\varphi_{i,j} = \varphi_{back}$  for the new time.

Unconditionally stable! (Why?)

(Though dissipative – loses energy over time.)

# Advection of Velocity

This handles scalars. What about advecting velocity?

$$\frac{\partial \mathbf{u}}{\partial t} = -\mathbf{u} \cdot \nabla \mathbf{u}$$

Same method:

- Trace back with current velocity
- Interpolate velocity at that point
- Assign it to the grid point at the *new time*.

Caution: Do not overwrite the velocity field you're using to trace back! (Make a copy.)

# 2. Pressure





# Recall... Helmholtz Decomposition

Input Velocity field

Curl-Free  
(irrotational)

Divergence-Free  
(incompressible)



$$u = \nabla p + \nabla \times \varphi$$

$$u_{old} = F_{pressure} + u_{new}$$

# Pressure Projection - Derivation

$$(1) \rho \frac{\partial \mathbf{u}}{\partial t} = -\nabla p \quad \text{and} \quad (2) \nabla \cdot \mathbf{u} = 0$$

Discretize (1) in time...

$$\mathbf{u}_{new} = \mathbf{u}_{old} - \frac{\Delta t}{\rho} \nabla p$$

Then plug into (2)...

$$\nabla \cdot \left( \mathbf{u}_{old} - \frac{\Delta t}{\rho} \nabla p \right) = 0$$

# Pressure Projection

Implementation:

1) Solve a linear system of equations for  $p$ :

$$\frac{\Delta t}{\rho} \nabla \cdot \nabla p = \nabla \cdot \mathbf{u}_{old}$$

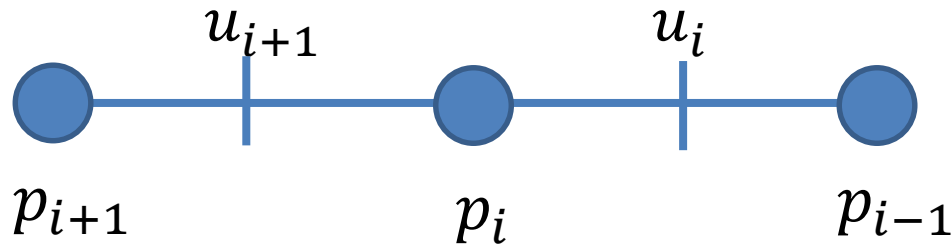
2) Given  $p$ , plug back in to update velocity:

$$\mathbf{u}_{new} = \mathbf{u}_{old} - \frac{\Delta t}{\rho} \nabla p$$

# Implementation

$$\frac{\Delta t}{\rho} \nabla \cdot \nabla p = \nabla \cdot \mathbf{u}_{old}$$

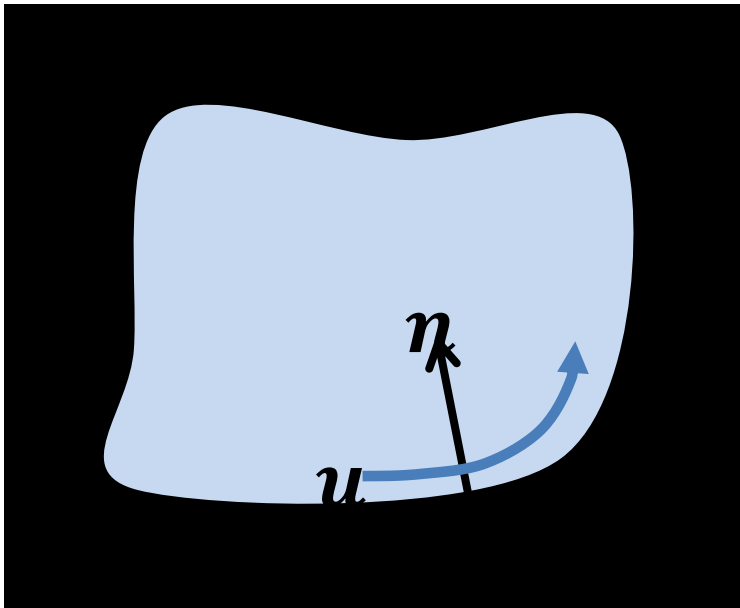
Discretize with finite differences (at staggered positions):



e.g., in 1D:

$$\frac{\Delta t}{\rho} \left( \frac{\frac{p_{i+1} - p_i}{\Delta x} - \frac{p_i - p_{i-1}}{\Delta x}}{\Delta x} \right) = \frac{u_{i+1}^{old} - u_i^{old}}{\Delta x}$$

# Solid Boundary Conditions



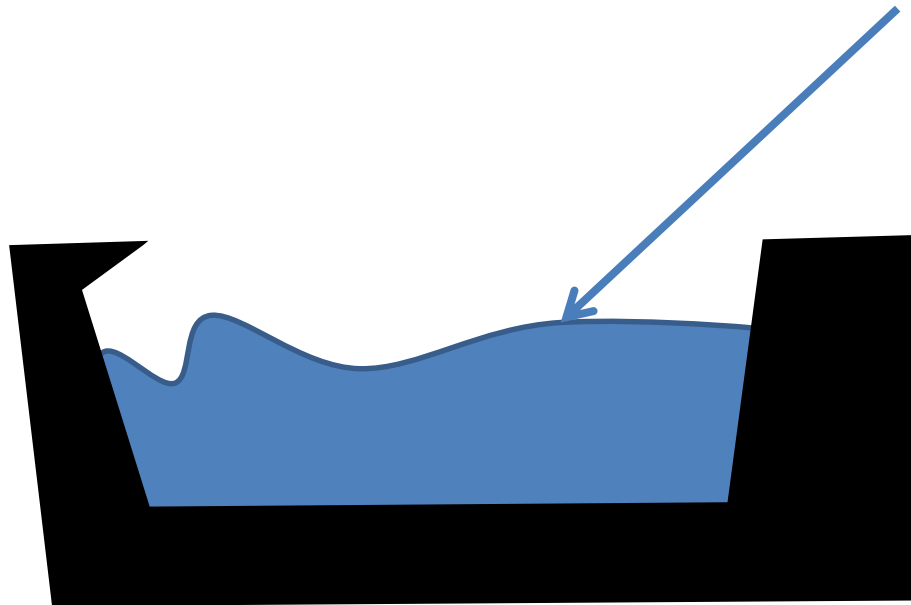
Free Slip:

$$\mathbf{u}_{new} \cdot \mathbf{n} = 0$$

i.e., Fluid cannot penetrate or flow out of the wall, but may slip along it.

# Air (“Free surface”) Boundary Conditions

Assume air (liquid exterior) is at some constant atmospheric pressure,  $p = p_{atm}$  or  $p = 0$ .



# 3. Viscosity

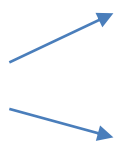


# Viscosity

$$\text{PDE: } \rho \frac{\partial \mathbf{u}}{\partial t} = \mu \nabla \cdot \nabla \mathbf{u}$$

Again, apply finite differences.

Discretized in time:

$$\mathbf{u}_{new} = \mathbf{u}_{old} + \frac{\Delta t \mu}{\rho} \nabla \cdot \nabla \mathbf{u}_*$$


$\mathbf{u}_{old}$  -> explicit time integration

$\mathbf{u}_{new}$  -> implicit time integration



# Viscosity – Time Integration

Explicit integration:  $\mathbf{u}_{new} = \mathbf{u}_{old} + \frac{\Delta t \mu}{\rho} \nabla \cdot \nabla \mathbf{u}_{old}$

- Compute  $\frac{\Delta t \mu}{\rho} \nabla \cdot \nabla \mathbf{u}_{old}$  from current velocities.
- Add on to current  $\mathbf{u}$ .
- Quite unstable (stability restriction:  $\Delta t \approx O(\Delta x^2)$ )

Implicit integration:  $\mathbf{u}_{new} = \mathbf{u}_{old} + \frac{\Delta t \mu}{\rho} \nabla \cdot \nabla \mathbf{u}_{new}$

- Stable even for high viscosities, large steps.
- Must solve a system of equations.

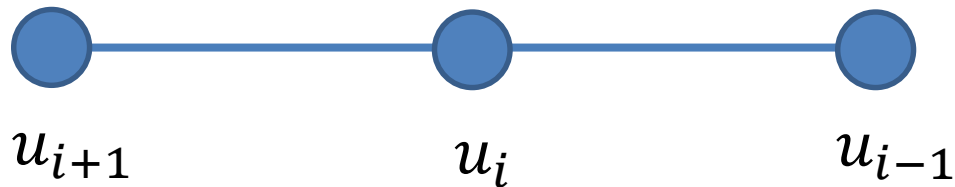
# Viscosity – Implicit Integration

Solve for  $\mathbf{u}_{new}$ :

$$\mathbf{u}_{new} - \frac{\Delta t \mu}{\rho} \nabla \cdot \nabla \mathbf{u}_{new} = \mathbf{u}_{old}$$

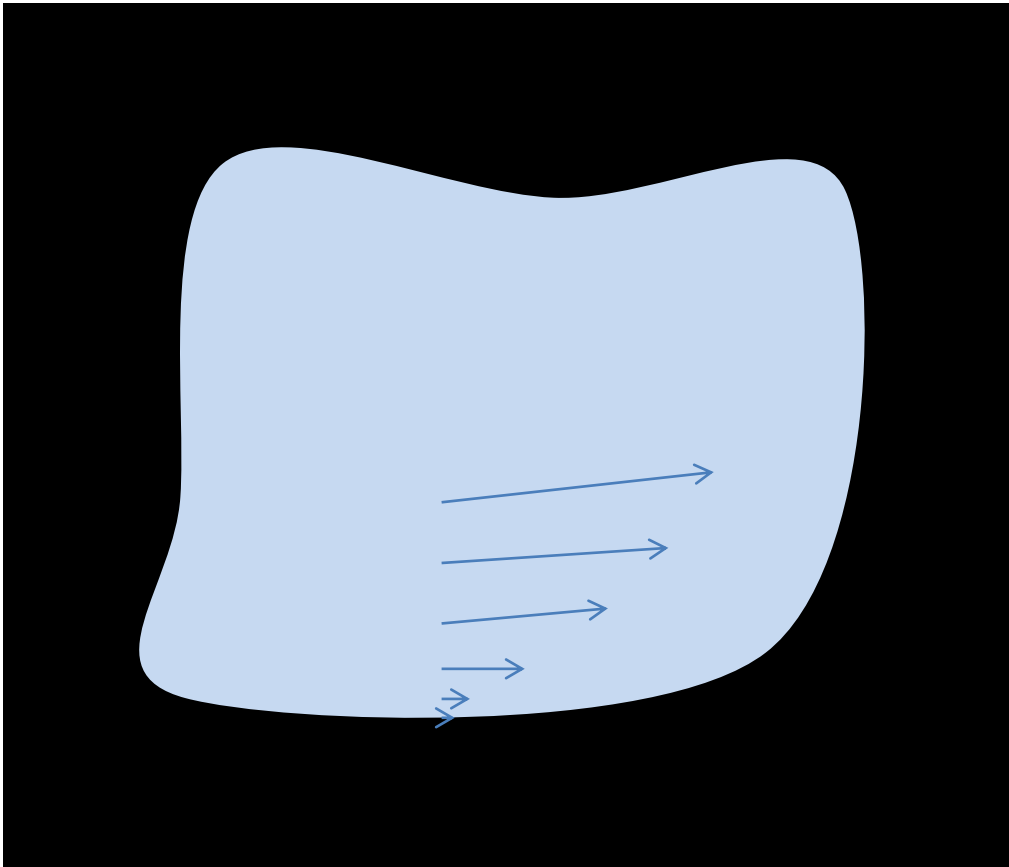
(Apply separately for each velocity component.)

e.g. in 1D:



$$u_i - \frac{\Delta t \mu}{\rho} \left( \frac{\frac{u_{i+1} - u_i}{\Delta x} - \frac{u_i - u_{i-1}}{\Delta x}}{\Delta x} \right) = u_i^{old}$$

# Viscosity - Solid Boundary Conditions



No-Slip:

$$\mathbf{u}_{new} = 0$$

# No-slip Condition

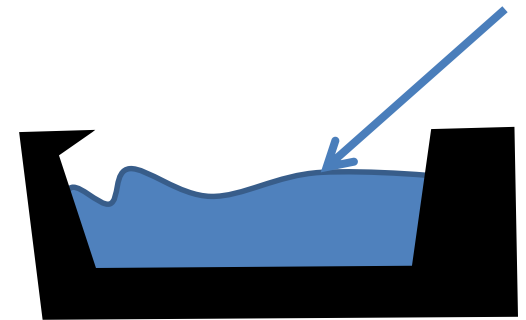


# Viscosity - Free Surface Conditions

Treat air as negligible, so enforce zero momentum exchange between liquid and “air”.

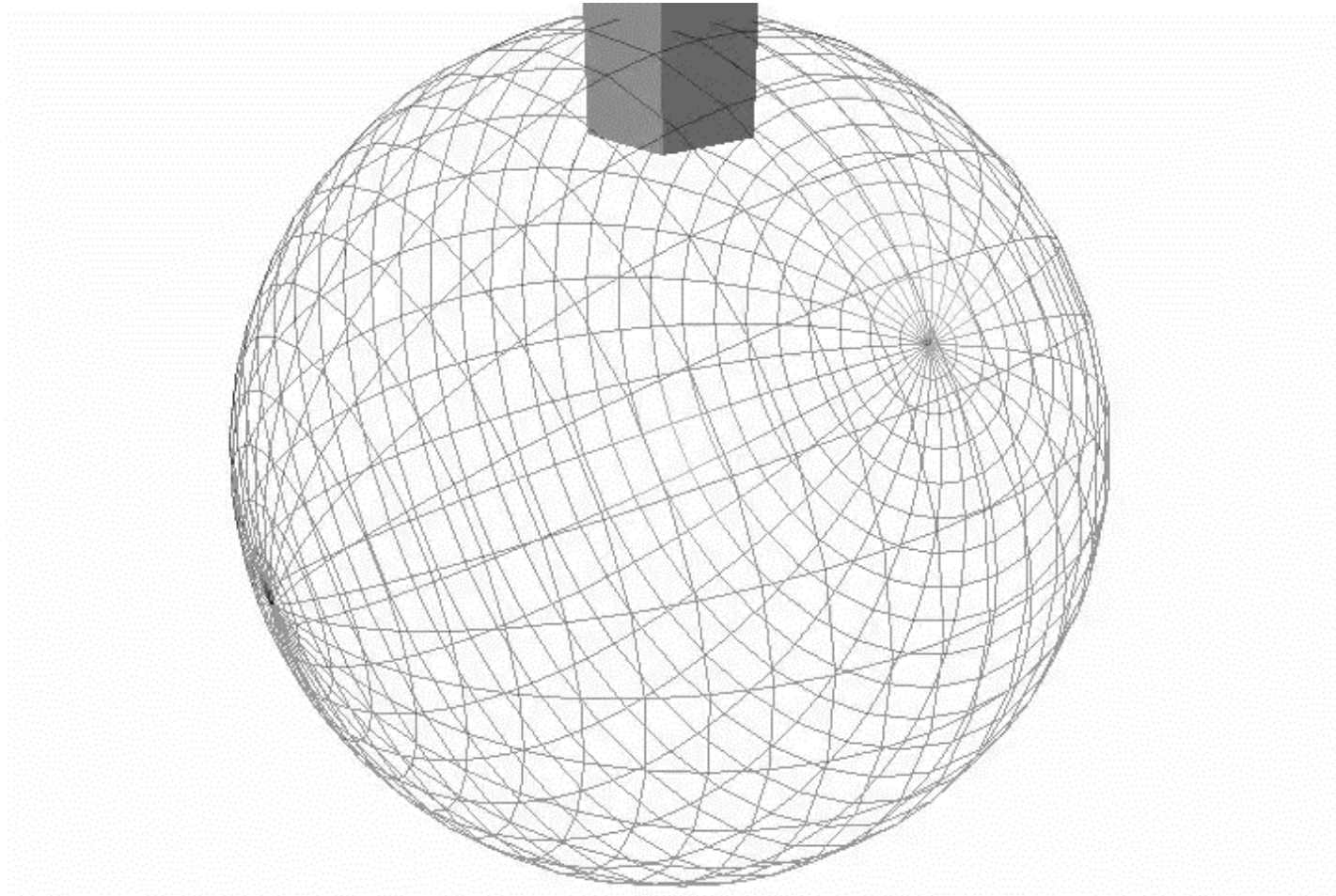
The proper conditions are quite involved:

$$\left( -p\mathbf{I} + \mu(\nabla\mathbf{u} + \nabla\mathbf{u}^T) \right) \cdot \mathbf{n} = \mathbf{0}$$



See [Batty & Bridson, 2008] for the standard solution in graphics. (Needed e.g., for honey coiling.)

# Viscous coiling simulation



# 4. External Forces



**Gravity.**

It's not just a good idea.  
It's the Law.



# Gravity

Discretized form is:

$$\mathbf{u}_{new} = \mathbf{u}_{old} + \Delta t \mathbf{g}$$

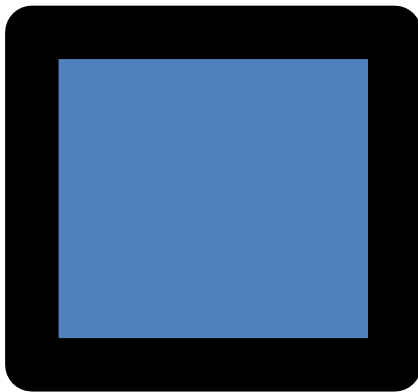
Simply increment the vertical velocities at each step!



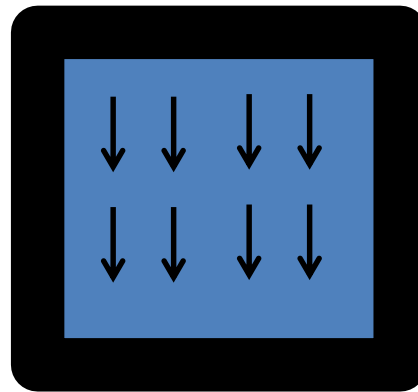
# Gravity

Notice: in a closed fluid-filled container, gravity (alone) won't do anything!

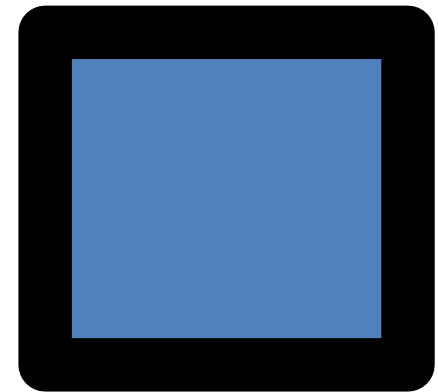
- Incompressibility cancels it out. (Assuming constant density.)



Start



After gravity step



After pressure step

# Simple Buoyancy

Track an extra scalar field  $T$ , representing local temperature, generated at a heat source.

Apply diffusion to  $T$ , and advect it along with the velocity field.

The difference between current and “reference” temperature induces a buoyancy force.



# Simple Buoyancy

e.g.

$$\mathbf{u}_{new} = \mathbf{u}_{old} + \Delta t \beta (T_{current} - T_{ref}) \mathbf{g}$$

$\beta$  dictates the strength of the buoyancy force.

For an enhanced version of this:

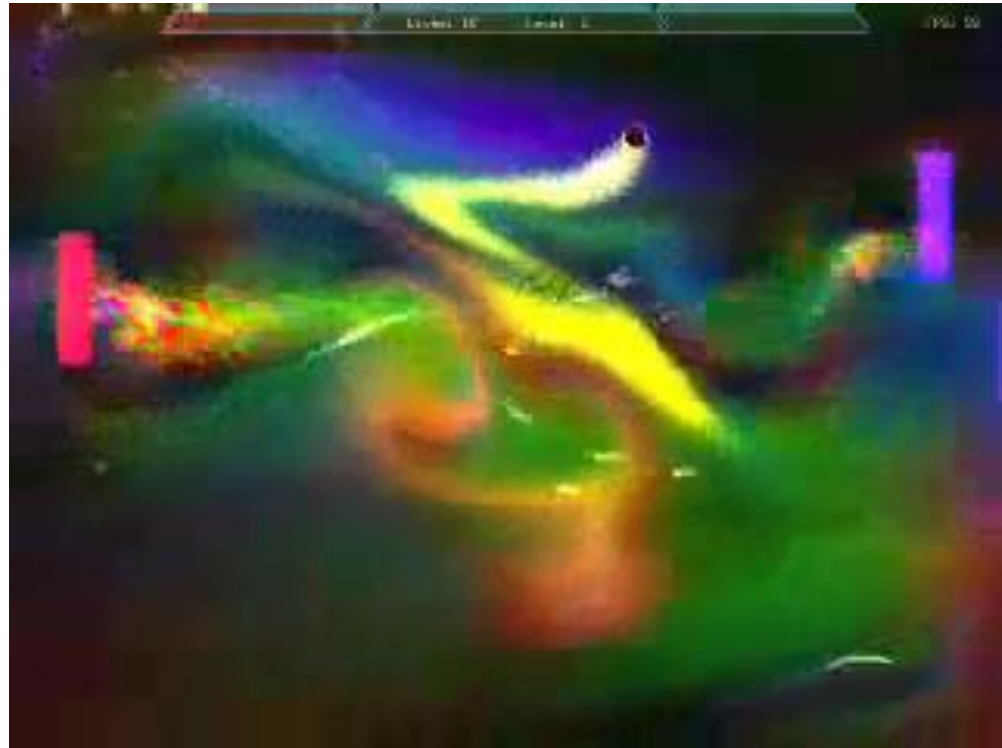
“Visual simulation of smoke”, [Stam et al., 2001].



# User Forces

Add whatever additional forces we want:

- Wind forces near a mouse click.
- Paddle forces in *Plasma Pong*.



*Plasma Pong game*

(eventually taken down due to copyright claim by Atari)

# Ordering of Steps

Order is important.

Why?

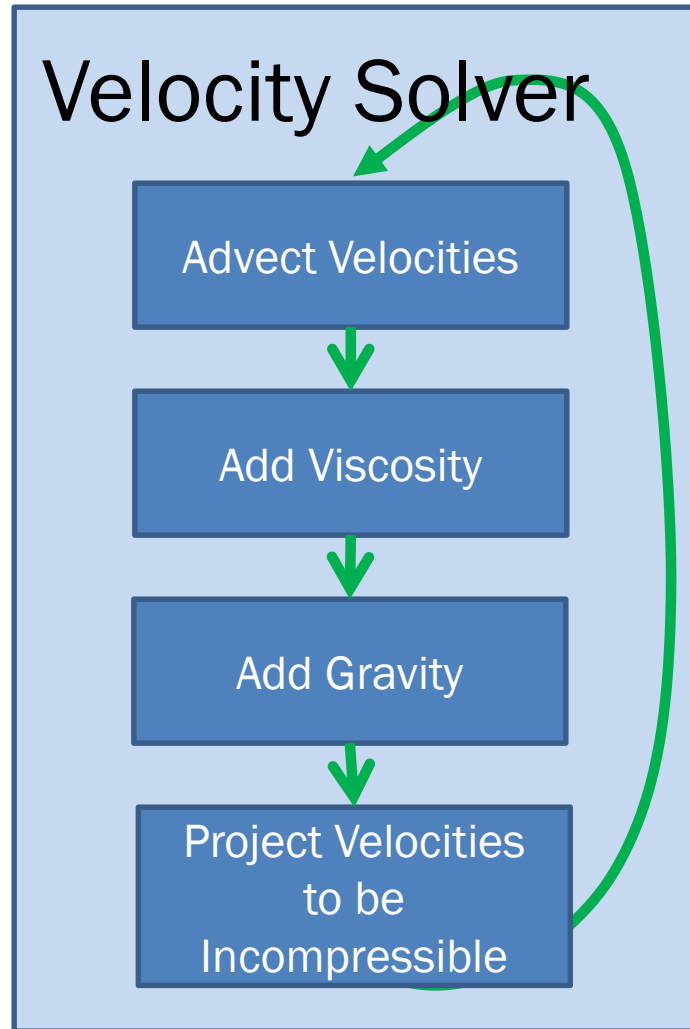
- 1) Incompressibility is not satisfied at intermediate steps.
- 2) Advecting with a divergent field causes volume/material loss or gain!

# Ordering of Steps

For example, consider advection in this field:



# The Big Picture



# Liquids





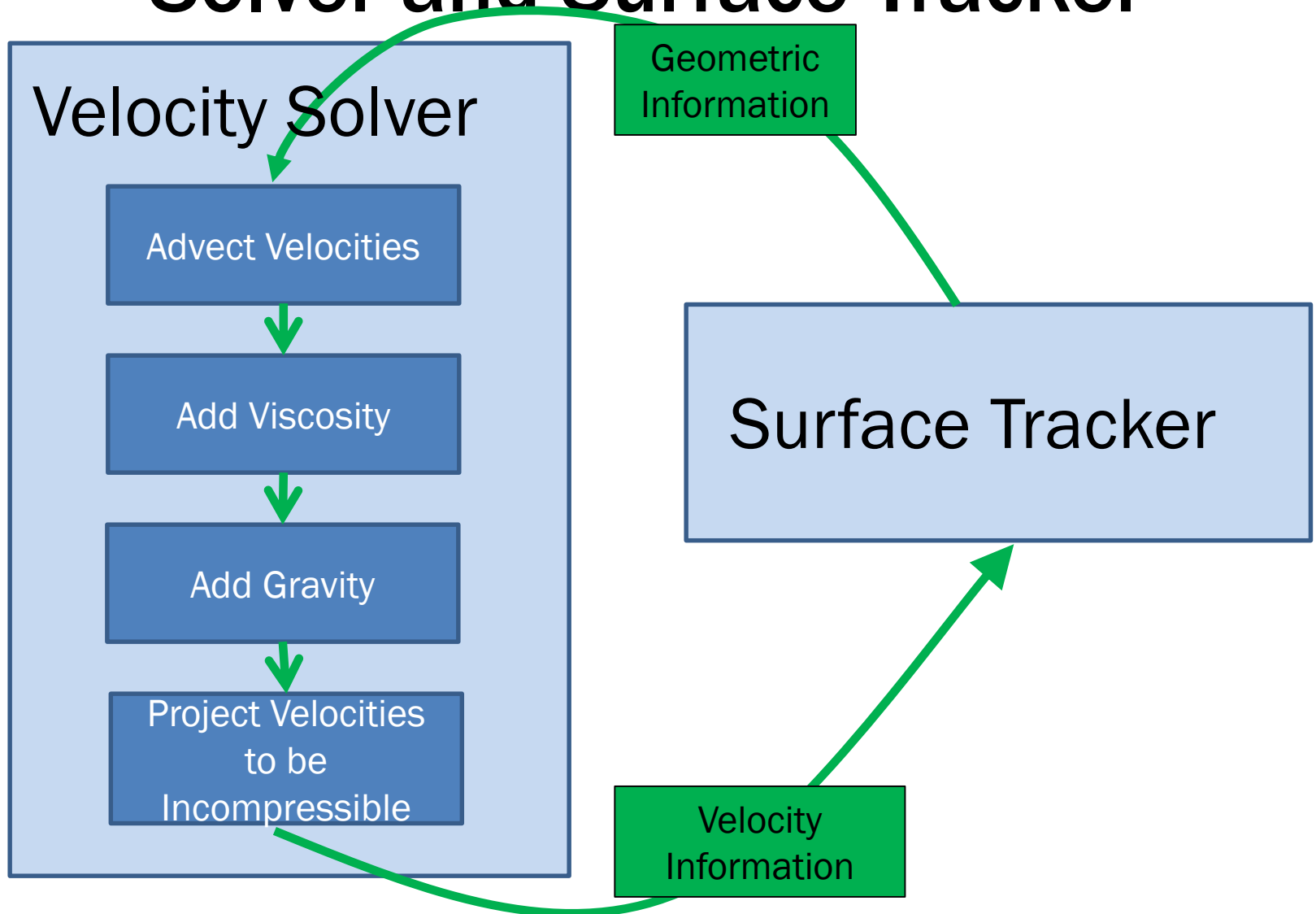
# Liquids

What's missing?



We still need a *surface representation*.

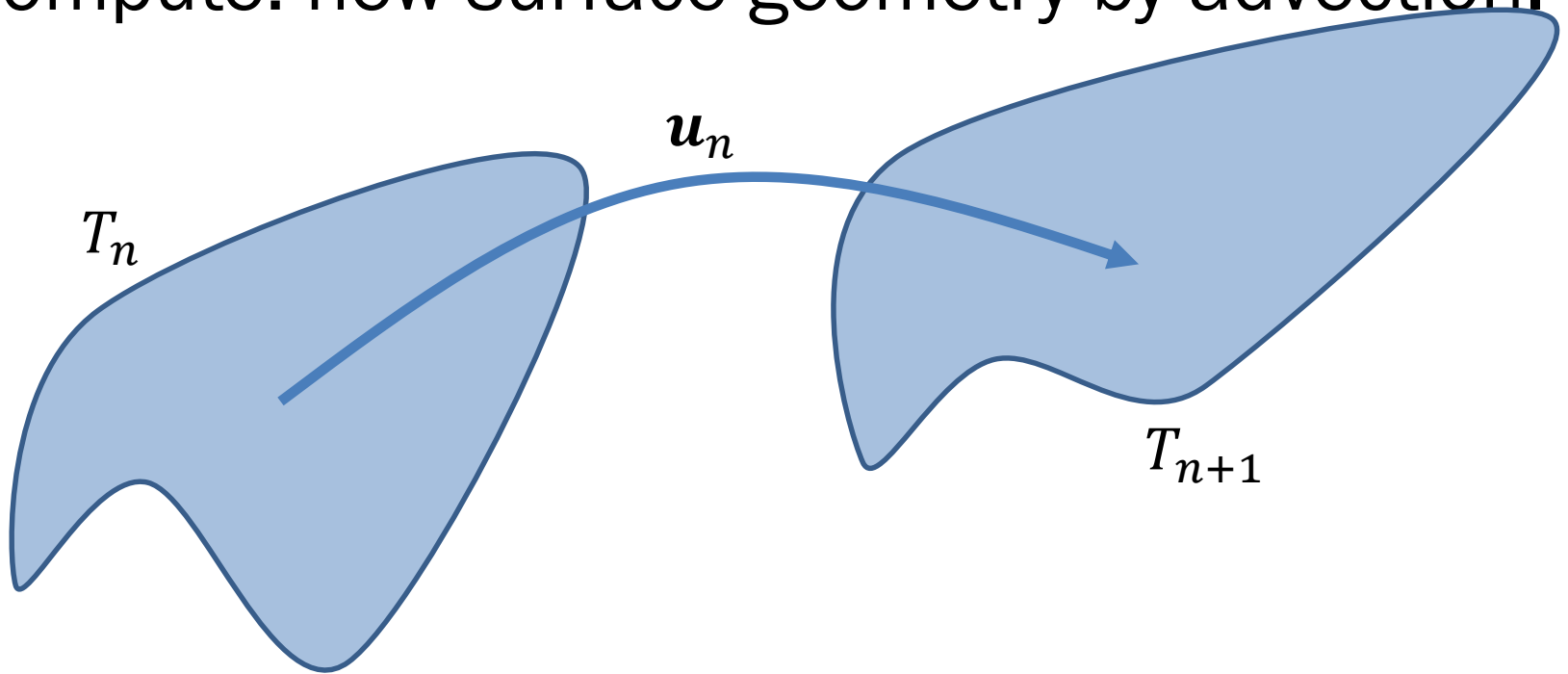
# Interaction between Solver and Surface Tracker



# Solver-to-Surface Tracker

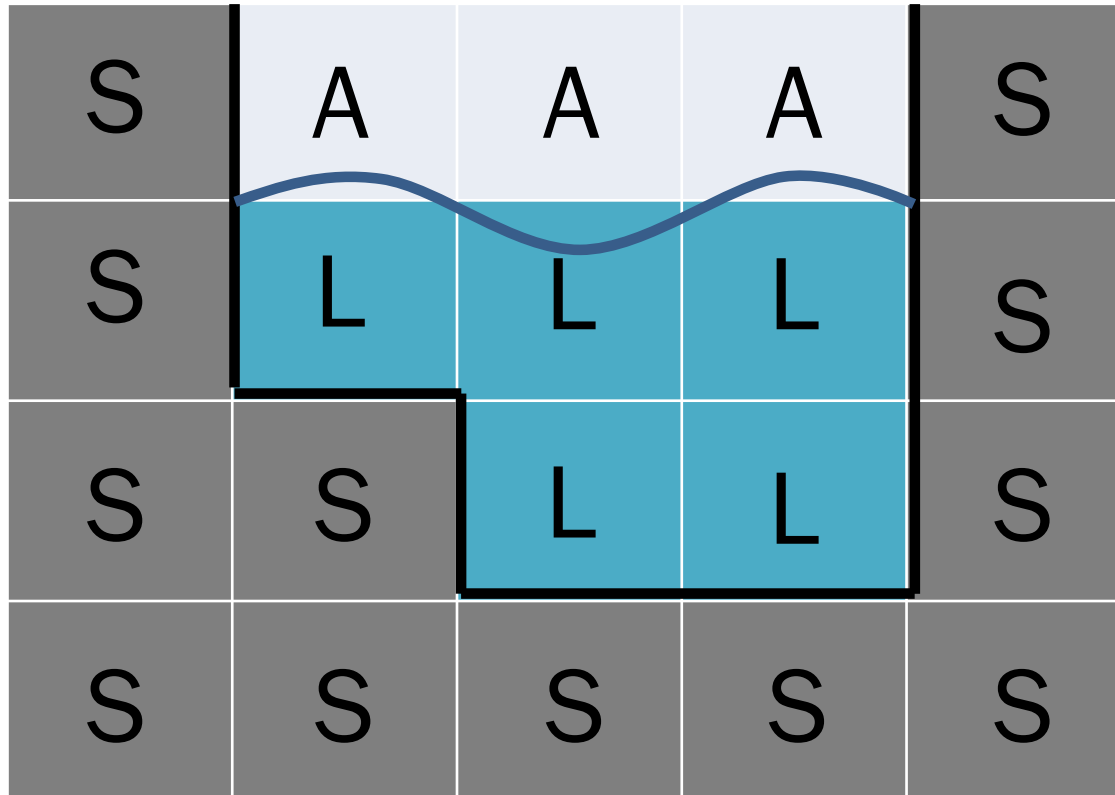
Given: current surface geometry, velocity field, and timestep.

Compute: new surface geometry by advection.



# Surface Tracker-to-Solver

Given the surface geometry, identify the type of each cell.  
Solver uses this information for boundary conditions.



# Surface Tracker

Ideally:

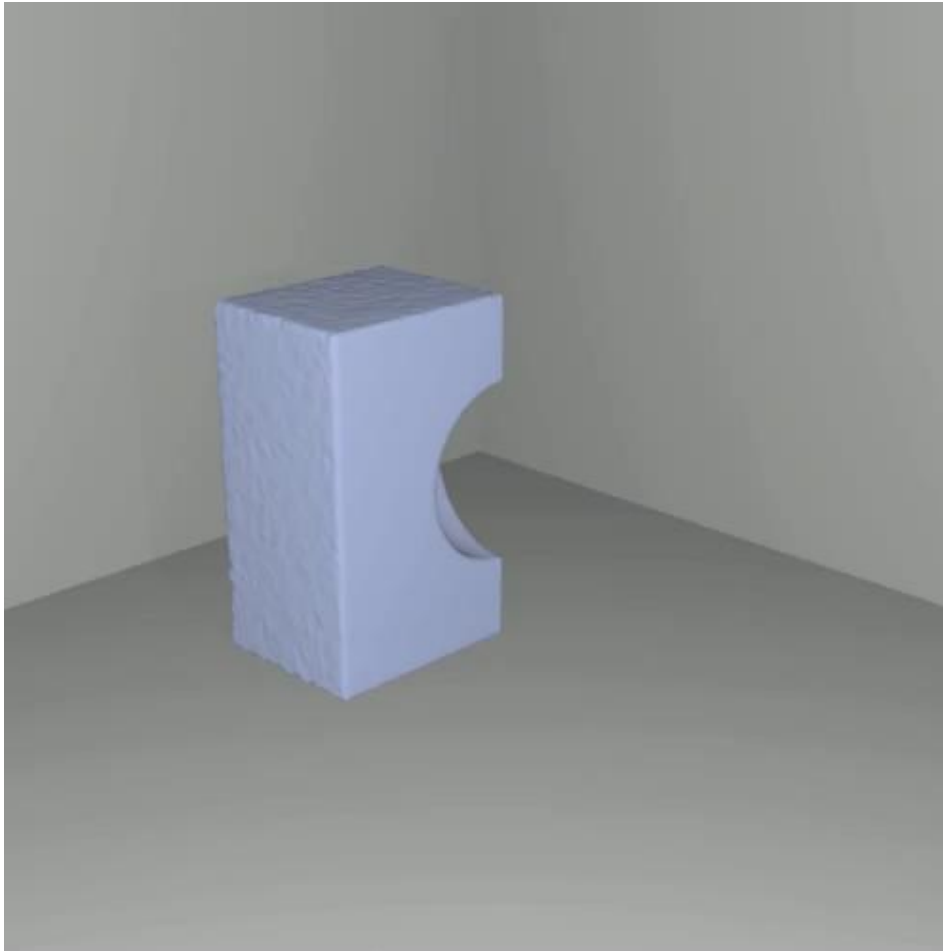
- Efficient
- Accurately follows the velocities
- Handles merging/splitting (“topology changes”)
- Conserves volume
- Retains small features and details
- Gives a smooth surface for rendering
- Provides convenient geometric operations (post-processing?)
- Easy to implement...

Very hard (impossible?) to do all of these at once.

# Surface Tracking Options

1. Particles
2. Level sets
3. Volume-of-fluid (VOF)
4. Triangle meshes
5. Hybrids (many of these)

# Particles

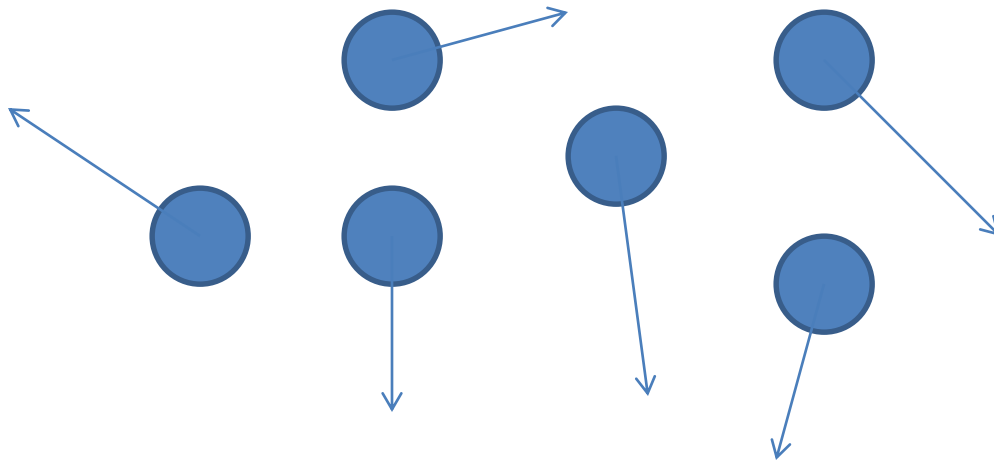


[Zhu & Bridson 2005]

# Particles

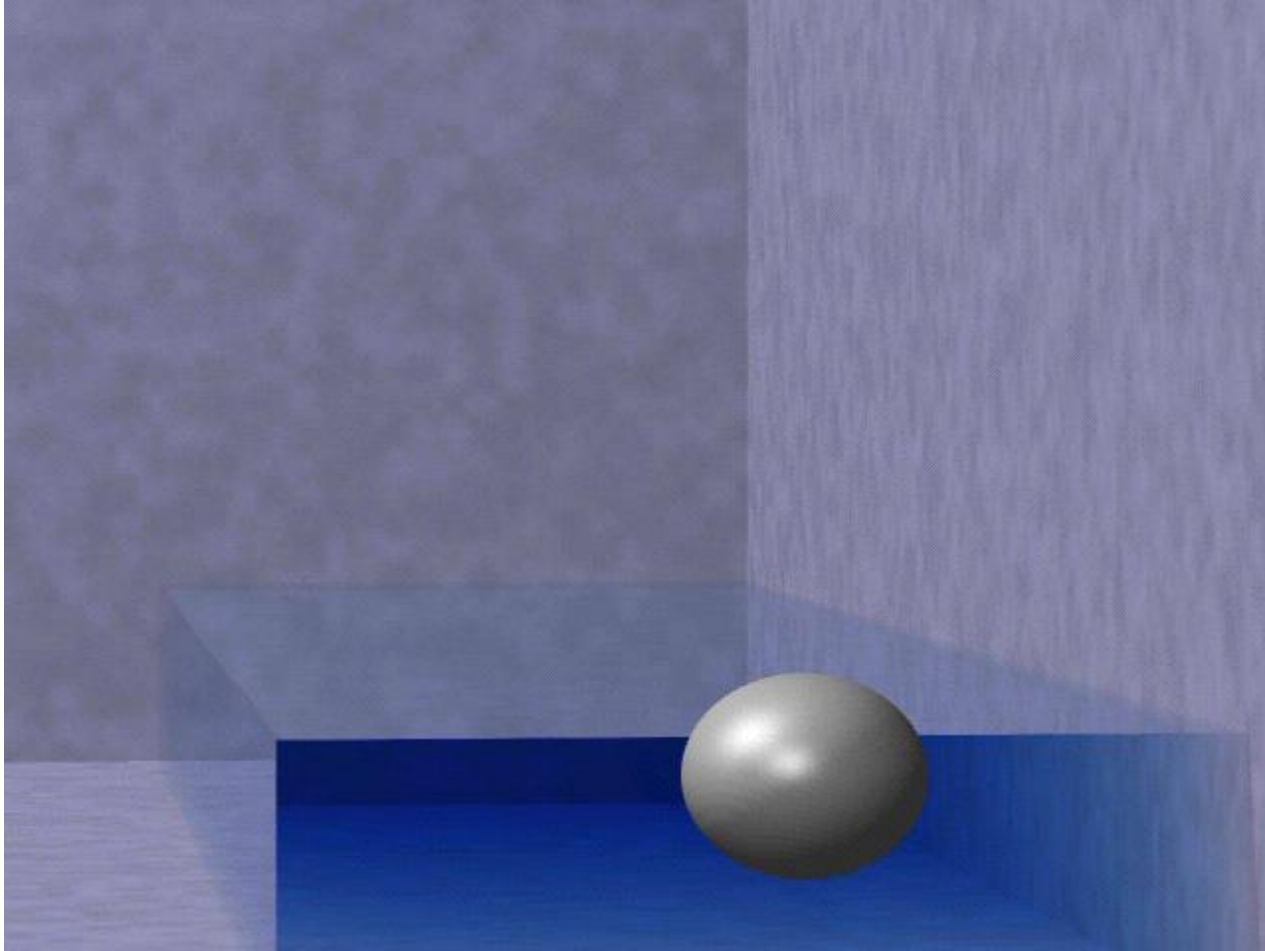
Perform passive Lagrangian advection on each particle.

For rendering, need to reconstruct a surface.





# Level sets

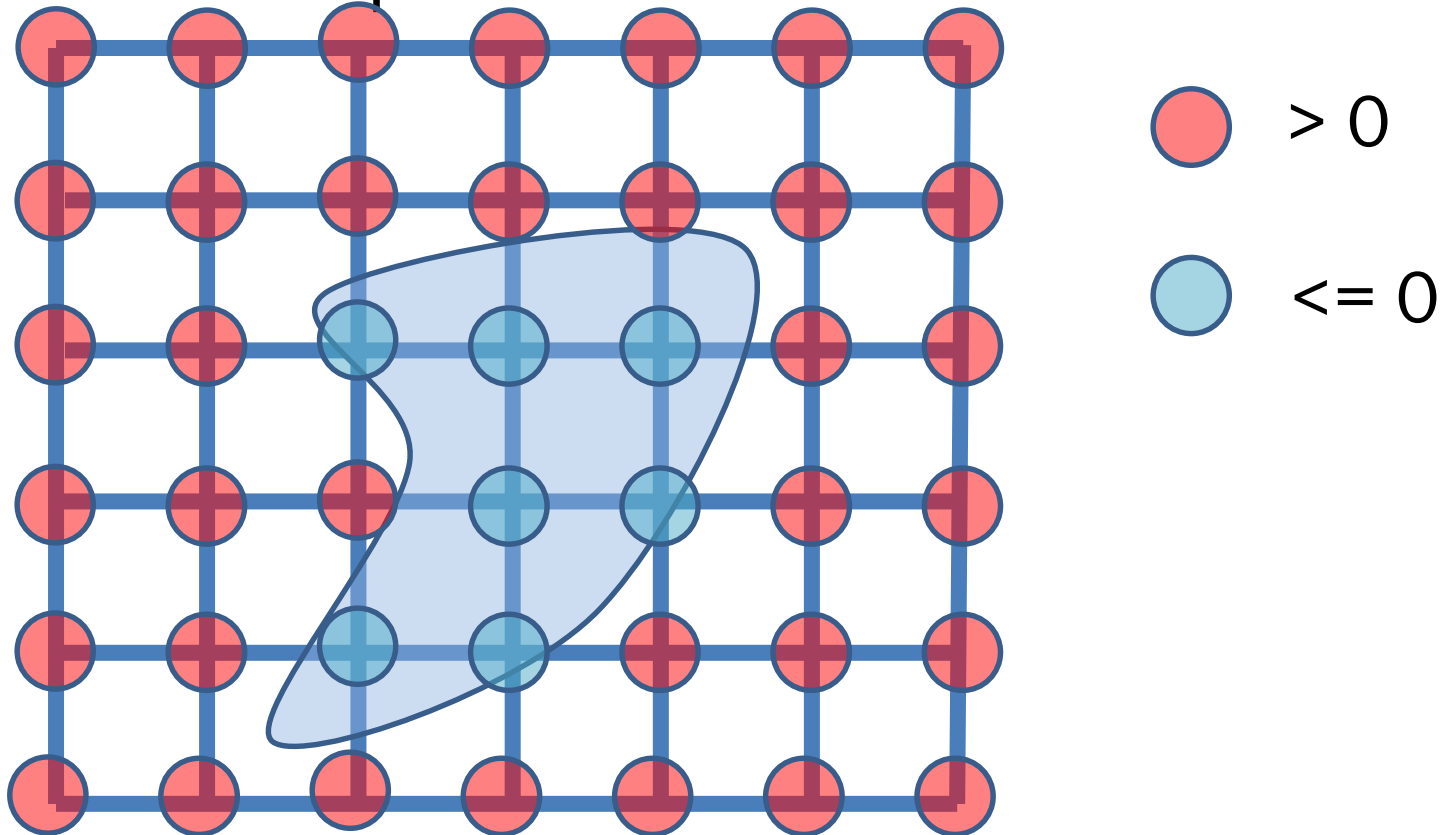


[Losasso et al.  
2004]

# Level sets

Each grid point stores *signed* distance to the surface (inside  $\leq 0$ , outside  $> 0$ ).

Surface is the interpolated zero isocontour.



# Densities / Volume of fluid

**Thin Surface Fluid Animation**

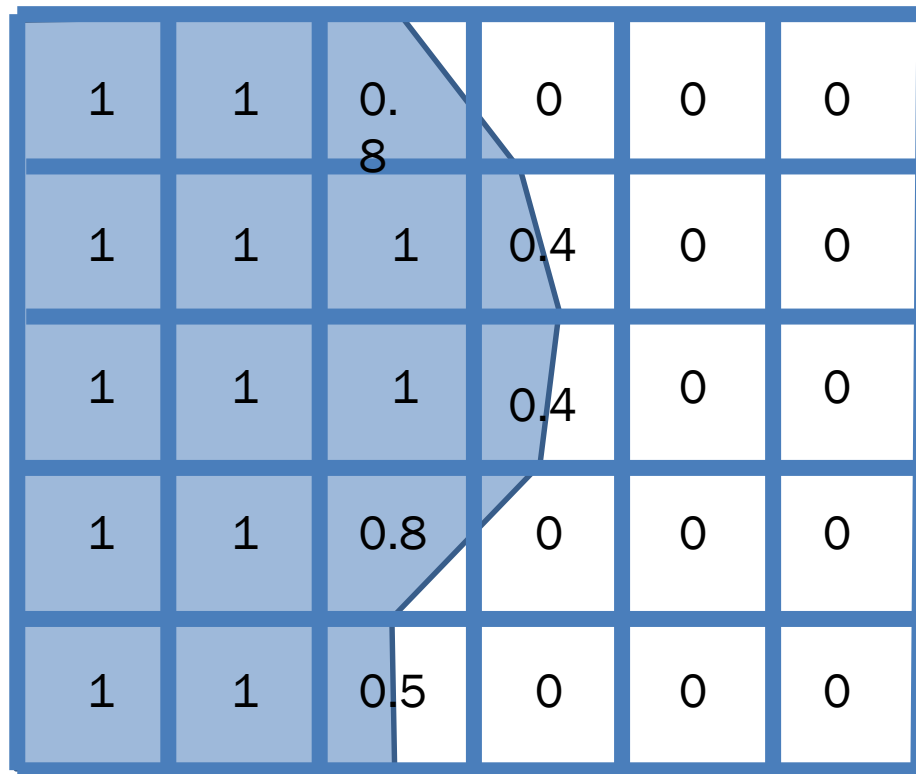
**Mass Density Resolution  $128^3$**

**Fluid Solver Resolution  $64^3$**

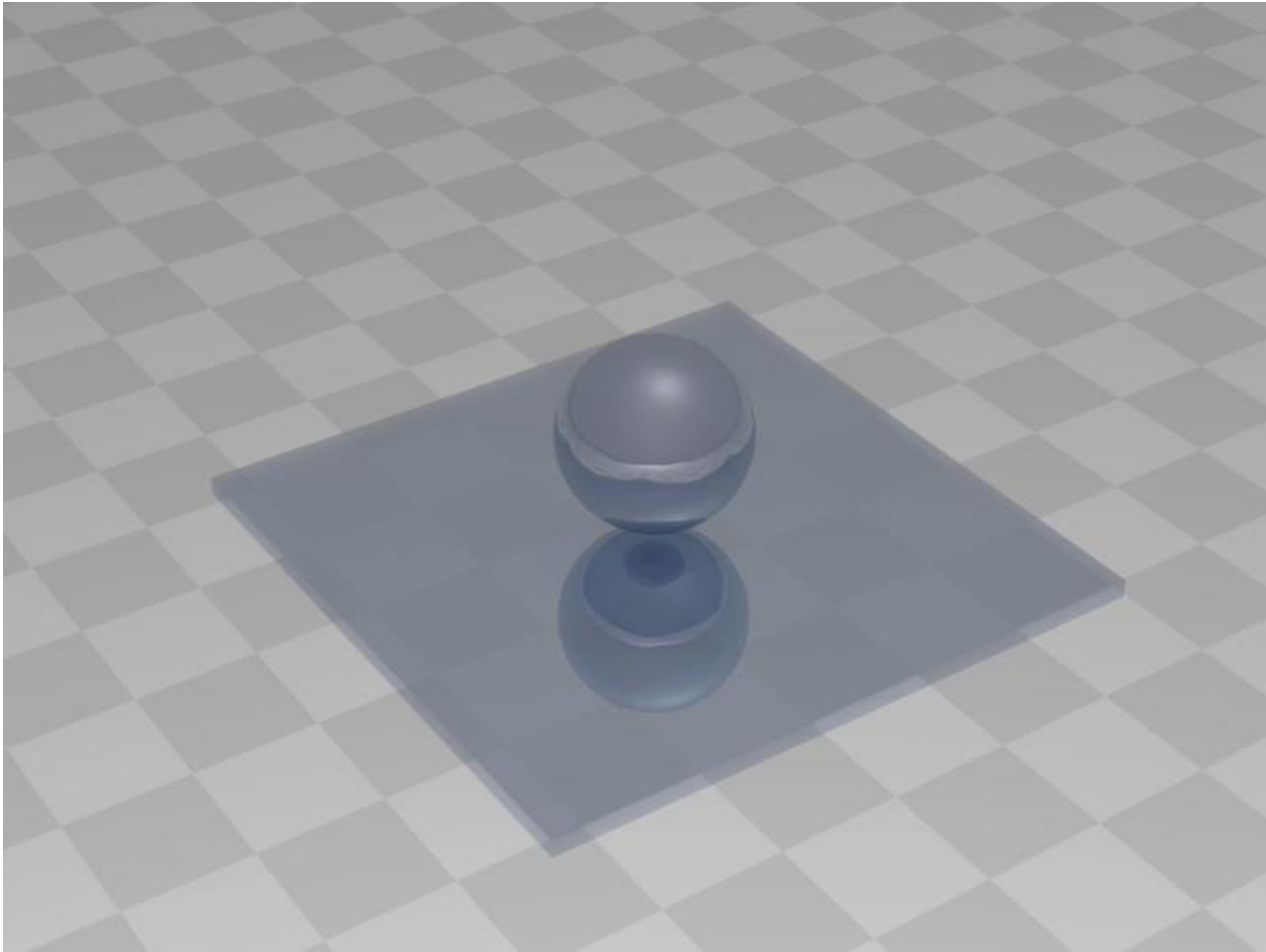
# Volume-Of-Fluid

Each cell stores fraction  $f \in [0,1]$  indicating how empty/full it is.

Surface is the transition region,  $f \approx 0.5$ .



# Meshes

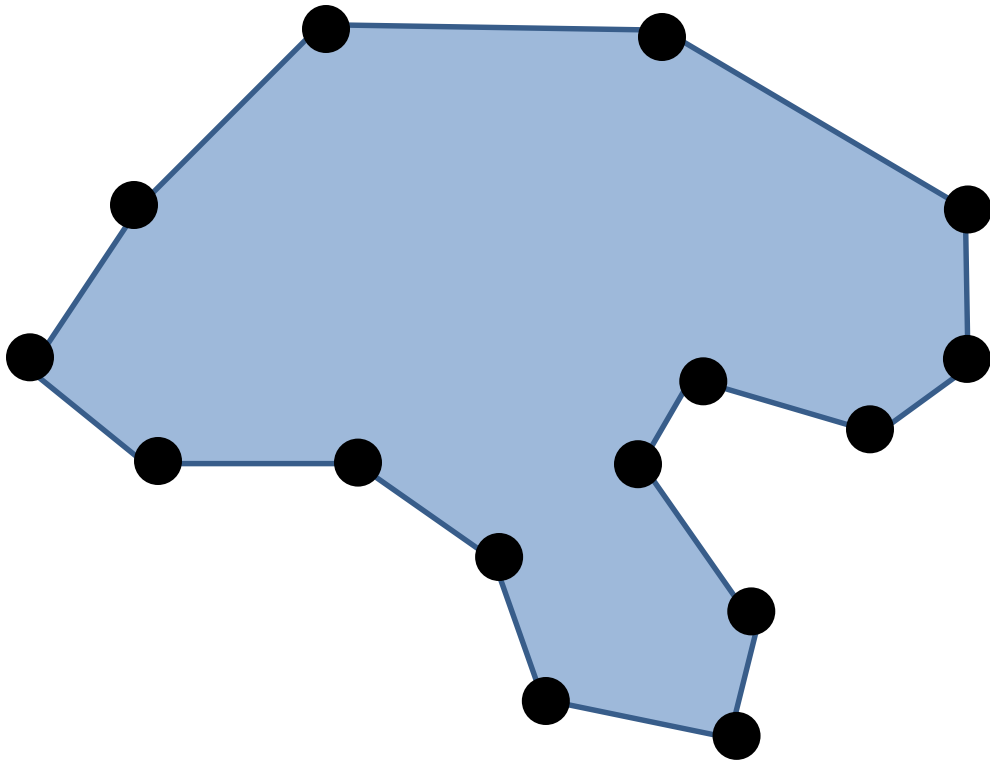


[Brochu et al 2010]

# Meshes

Store a triangle mesh.

Advect its vertices, and deal with collisions.



# Reminders

- 1<sup>st</sup> round presentations start Monday.  
Graded on...
  1. Knowledge/coverage of technical concepts
  2. Organization
  3. Slide quality
  4. Speaking/presentation skills
- 1<sup>st</sup> round of paper reviews due Sunday, 5pm.
- Start thinking about project topics.
- Piazza (or email) with any questions.

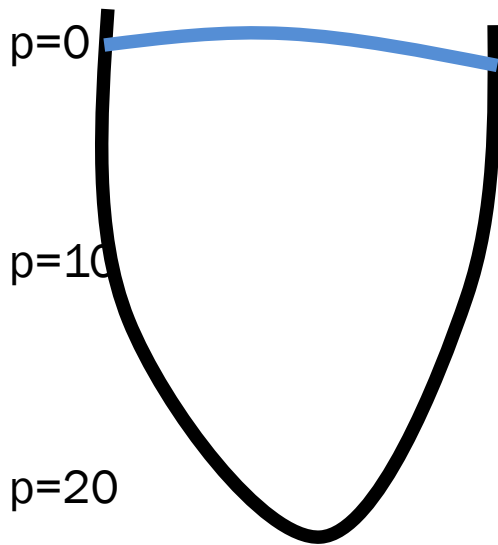






# Free Surface Boundary Conditions

Only the pressure *gradient* matters, so simplify by assuming  $p = p_{atm} = 0$ .



Same (vertical)  
pressure  
gradient,  $\nabla p$ .

